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# Naturalness Bounds on Neutrino Magnetic Moments

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# Outline

- Introduction: basic facts on neutrino magnetic moments
  - “Naturalness bounds”
    - Dirac neutrino magnetic moments
    - Majorana neutrino transition moments
- } Quite different bounds !!
- Understanding the difference on the basis of symmetry

- N.F. Bell, VC, M.J. Ramsey-Musolf, P.Vogel, M. Wise, PRL 95 151802, 2005
- S. Davidson, M Gorbahn, A. Santamaria PLB 626 151, 2005
- N.F. Bell, M. Gorchtein, M.J. Ramsey-Musolf, P.Vogel, P.Wang, PLB 642 377, 2006

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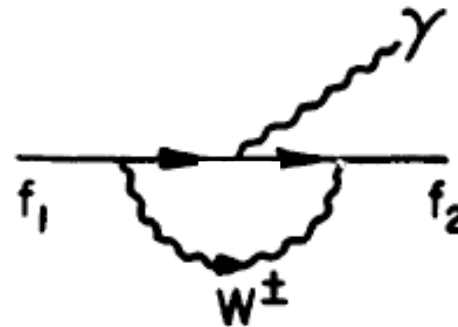
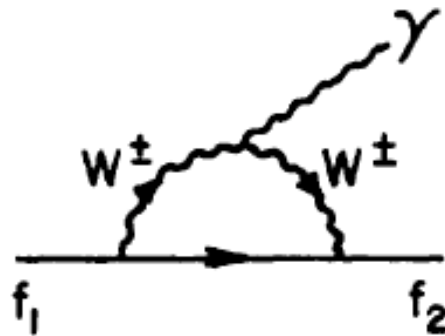
# Neutrino magnetic moments: facts

- Massive neutrinos have non-zero magnetic moment.  
In minimally extended SM

$$\frac{\mu_\nu}{\mu_B} = \frac{3 G_F m_e m_\nu}{4\sqrt{2} \pi^2} \approx 3 \times 10^{-19} \left[ \frac{m_\nu}{1 \text{ eV}} \right]$$

$$\mu_B = e/(2m_e)$$

Lee-Shrock '77  
Marciano-Sanda '77



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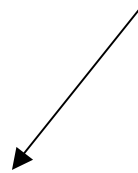
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Lee-Shrock '77  
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$$\mu_B = e/(2m_e)$$

- Most stringent **phenomenological bounds** come from:

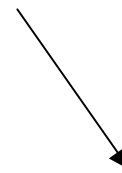
$\nu e \rightarrow \nu e$



$$\frac{|\mu_\nu^{\text{exp}}|}{\mu_B} \simeq \frac{G_F m_e}{\sqrt{2}\pi\alpha} \sqrt{m_e T_{\text{kin}}} \sim 10^{-10}$$

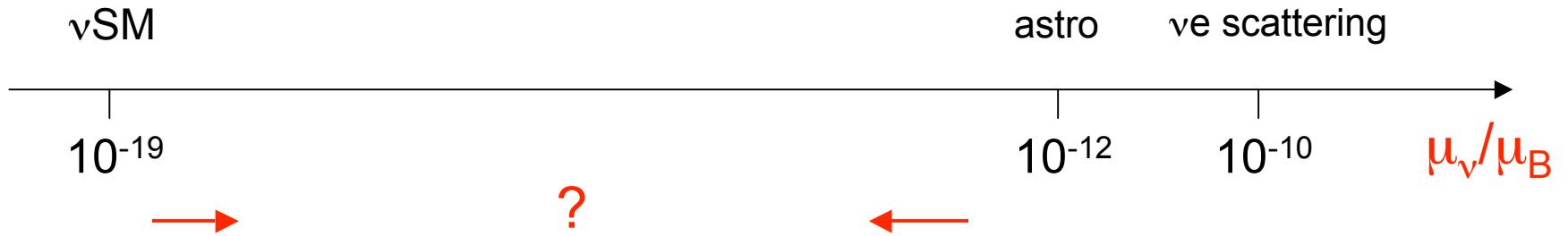
Energy loss in stars via

$\gamma^* \rightarrow \nu \bar{\nu}$



$$\frac{|\mu_\nu^{\text{astro}}|}{\mu_B} \simeq \frac{G_F m_e}{\sqrt{2}\pi\alpha} \omega_P \sim 3 \times 10^{-12}$$

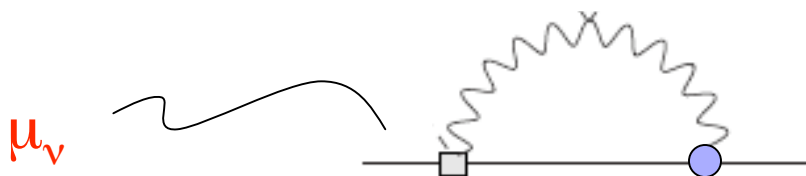
# Theoretical considerations



*How large a magnetic moment can we expect ?*

# Theoretical considerations

- We expect BSM physics to generate  $\mu_\nu/\mu_B > 10^{-19}$
- Several models can make  $\mu_\nu/\mu_B \gg 10^{-19}$  despite  $m_\nu < 1 \text{ eV}$   
(based on *symmetries that are broken by SM interactions*)
- SM radiative corrections induce  $\delta m_\nu \propto \mu_\nu \Rightarrow$  get  
“naturalness” upper bounds on  $\mu_\nu$  by requiring  $\delta m_\nu < m_\nu^{\text{obs}}$





## Naturalness bounds: Dirac $\nu$

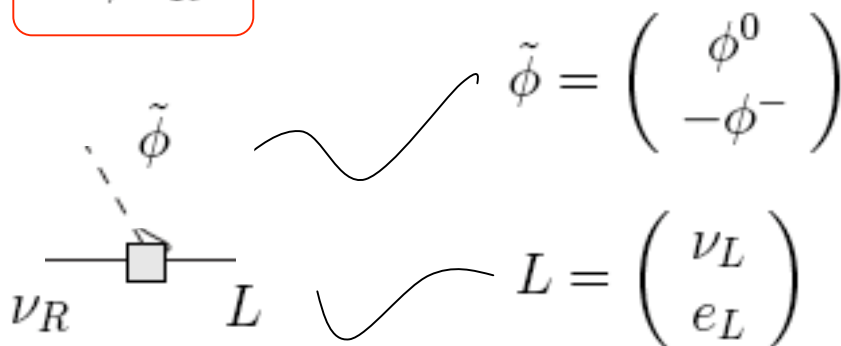
- Assume new physics responsible for  $\mu_\nu$  at scale  $\Lambda > v_{ew}$
- Work within EFT at  $E < \Lambda$  that respects  $SU(2) \times U(1)$  and L

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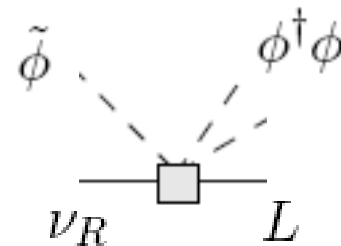
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$$\mathcal{L}_{\text{eff}} \supset C_M^{(4)} \mathcal{O}_M^{(4)} + \frac{C_M^{(6)}}{\Lambda^2} \mathcal{O}_M^{(6)} + \frac{C_B}{\Lambda^2} \mathcal{O}_B^{(6)} + \frac{C_W}{\Lambda^2} \mathcal{O}_W^{(6)}$$

$$\bar{L} \tilde{\phi} \nu_R$$



$$\bar{L} \tilde{\phi} \nu_R (\phi^\dagger \phi)$$

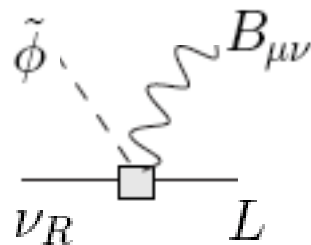


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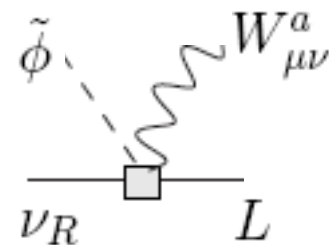
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$$g_1 \bar{L} \tilde{\phi} \sigma^{\mu\nu} \nu_R B_{\mu\nu}$$



$$g_2 \bar{L} \tau^a \tilde{\phi} \sigma^{\mu\nu} \nu_R W_{\mu\nu}^a$$



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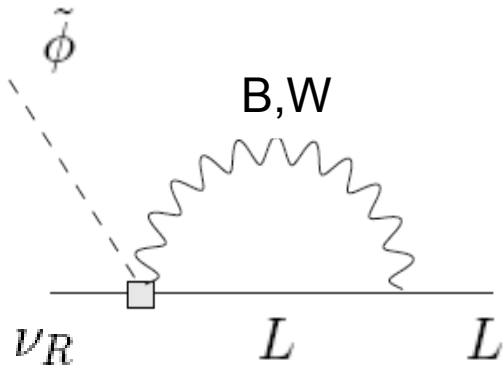
$$\mathcal{L}_{\text{eff}} \supset C_M^{(4)} \mathcal{O}_M^{(4)} + \frac{C_M^{(6)}}{\Lambda^2} \mathcal{O}_M^{(6)} + \frac{C_B}{\Lambda^2} \mathcal{O}_B^{(6)} + \frac{C_W}{\Lambda^2} \mathcal{O}_W^{(6)}$$

$$\phi \rightarrow \begin{pmatrix} 0 \\ v/\sqrt{2} \end{pmatrix}$$

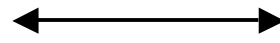
$$m_\nu = -C_M^{(4)}(v) \frac{v}{\sqrt{2}} - C_M^{(6)}(v) \frac{v^3}{2\sqrt{2}\Lambda^2}$$

$$\frac{\mu_\nu}{\mu_B} = -4\sqrt{2} \left( \frac{m_e v}{\Lambda^2} \right) [C_B(v) + C_W(v)]$$

- $\mathcal{O}_{B,W}^{(6)} \rightarrow \mathcal{O}_M^{(4)}$  “mixing”

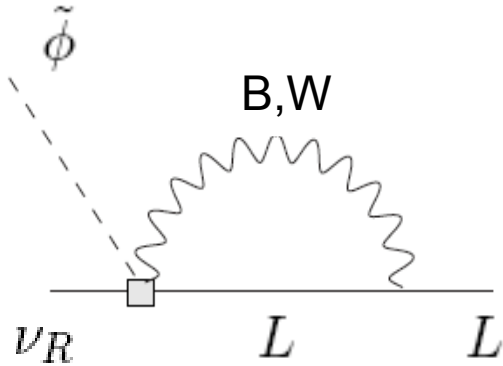


$$\delta C_M^{(4)}(\Lambda) \sim \frac{\alpha_{1,2}}{4\pi} C_{B,W}(\Lambda)$$

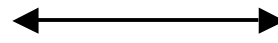


$$\delta m_\nu \sim \frac{\alpha}{4\pi} \frac{\Lambda^2}{m_e} \frac{\mu_\nu}{\mu_B}$$

- $\mathcal{O}_{B,W}^{(6)} \rightarrow \mathcal{O}_M^{(4)}$  “mixing” (*matching* at scale  $\Lambda$ )



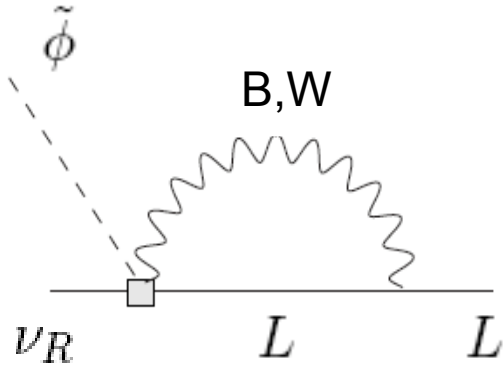
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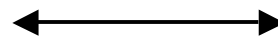
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coefficient is model-dependent

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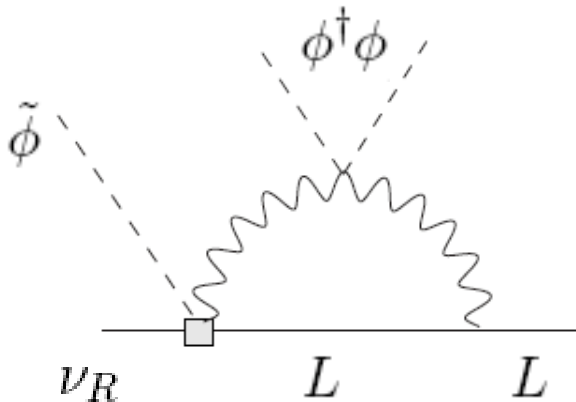
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coefficient is model-dependent

- $\mathcal{O}_{B,W}^{(6)} \rightarrow \mathcal{O}_M^{(6)}$  *mixing*



$$\delta m_\nu \sim \frac{\alpha}{4\pi} \frac{v^2}{m_e} \left( \alpha \log \frac{\Lambda}{M_Z} \right) \frac{\mu_\nu}{\mu_B}$$

+ ...

coefficient is model independent

- Resulting bounds on Dirac  $\nu$  magnetic moment:

Dim-4  
“matching”

$$\frac{|\mu_\nu|}{\mu_B} \lesssim 3 \times 10^{-15} \times \left( \frac{\delta m_\nu}{1 \text{ eV}} \right) \times \left( \frac{1 \text{ TeV}}{\Lambda} \right)^2$$

Strong dependence on  $\Lambda$

Dim-6  
“mixing”

$$\frac{|\mu_\nu|}{\mu_B} \lesssim 8 \times 10^{-15} \times \left( \frac{\delta m_\nu}{1 \text{ eV}} \right)$$

Mild logarithmic dependence on  $\Lambda$

At  $\Lambda \sim 1 \text{ TeV}$  both bounds are  $10^2$  ( $10^4$ ) stronger than astrophysical (laboratory) limits

# Naturalness bounds: Majorana $\nu$

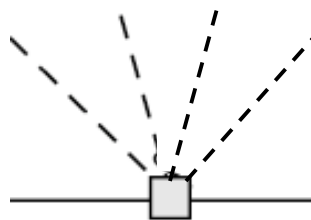
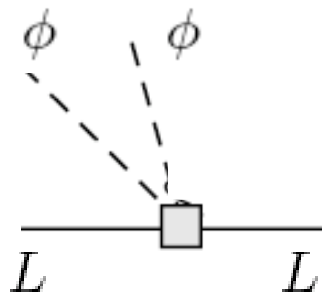
- SU(2) x U(1) invariant EFT at  $E < \Lambda$  [but L is not conserved]

$$\mathcal{L}_{\text{eff}} \supset \frac{C_M^{(5)}}{\Lambda} \mathcal{O}_M^{(5)} + \frac{C_M^{(7)}}{\Lambda^3} \mathcal{O}_M^{(7)} + \frac{C_B}{\Lambda^3} \mathcal{O}_B^{(7)} + \frac{C_W}{\Lambda^3} \mathcal{O}_W^{(7)}$$

$$(\overline{L}_\alpha^c \epsilon \phi) (\phi^T \epsilon L_\beta)$$



$$(\overline{L}_\alpha^c \epsilon \phi) (\phi^T \epsilon L_\beta) (\phi^\dagger \phi)$$



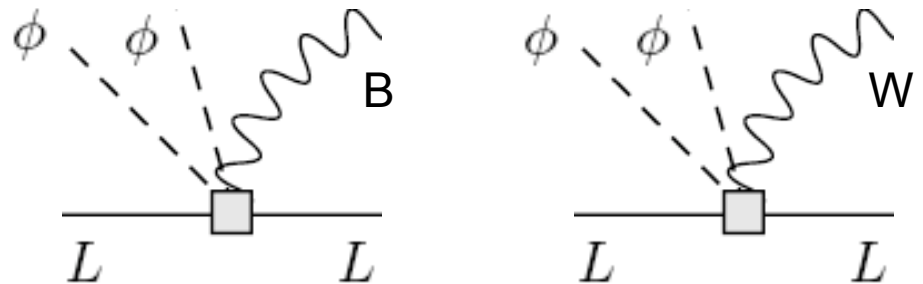
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$$g_1 (\bar{L}_\alpha^c \epsilon \phi) \sigma^{\mu\nu} (\phi^T \epsilon L_\beta) B_{\mu\nu}$$

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$$[m_\nu]_{\alpha\beta} = [m_\nu]_{\beta\alpha}$$

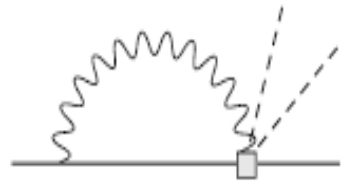
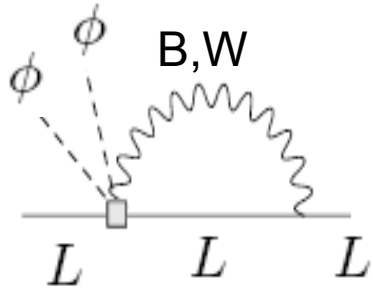
$$\phi \rightarrow \begin{pmatrix} 0 \\ v/\sqrt{2} \end{pmatrix}$$

$$\frac{1}{2} [m_\nu]_{\alpha\beta} = \frac{v^2}{2\Lambda} [C_M^{(5)}(v)]_{\alpha\beta} + \frac{v^4}{4\Lambda^3} [C_M^{(7)}(v)]_{\alpha\beta}$$

$$[\mu_\nu]_{\alpha\beta} = -[\mu_\nu]_{\beta\alpha}$$

$$\frac{[\mu_\nu]_{\alpha\beta}}{\mu_B} = \frac{2m_e v^2}{\Lambda^3} \left( [C_B(v)]_{\alpha\beta} + [C_W(v)]_{\alpha\beta} \right)$$

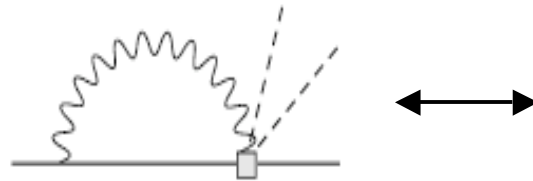
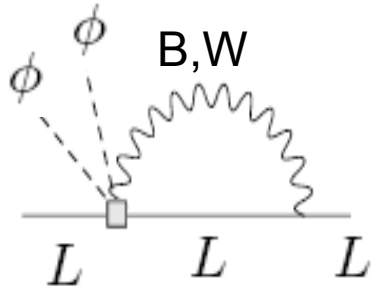
- $\mathcal{O}_{B,W}^{(7)} \rightarrow \mathcal{O}_M^{(5)}$  “mixing” (*matching* at scale  $\Lambda$ )



The sum vanishes, as expected !

$$m^T = m \quad \mu^T = -\mu$$

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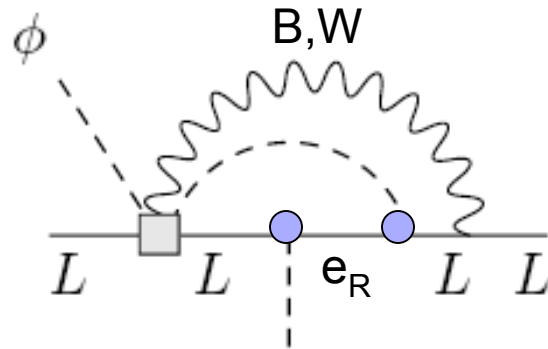
$$m^T = m \quad \mu^T = -\mu$$

Need to insert non-trivial flavor structure:  
provided by charged lepton Yukawa couplings

$$Y_\alpha = \frac{m_\alpha}{v}$$

A Feynman diagram illustrating a charged lepton Yukawa coupling. A horizontal line has a vertex labeled  $L$  on the left and  $e_R$  on the right. A wavy line connects the vertex to the left. A dashed vertical line labeled  $\phi$  extends upwards from the vertex.

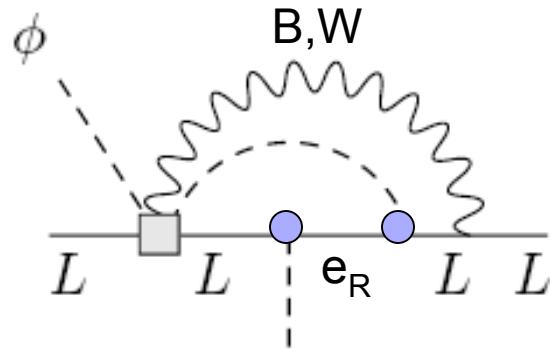
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$$\delta[m_\nu]_{\alpha\beta} \sim \frac{\alpha}{4\pi} \frac{\Lambda^2}{m_e} \frac{[\mu_\nu]_{\alpha\beta}}{\mu_B} \cdot \left[ \frac{Y_\alpha^2 - Y_\beta^2}{(4\pi)^2} \right]$$

coefficient is model-dependent

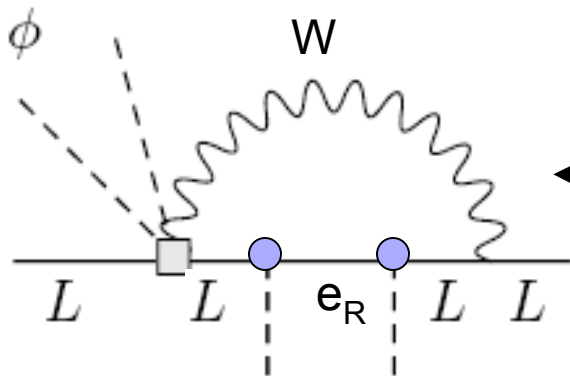
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- $\mathcal{O}_{B,W}^{(7)} \rightarrow \mathcal{O}_M^{(7)}$  *mixing*



$$\delta[m_\nu]_{\alpha\beta} \sim \frac{\alpha}{4\pi} \frac{v^2}{m_e} \log \frac{\Lambda}{M_Z} \frac{[\mu_\nu]_{\alpha\beta}}{\mu_B} \cdot [Y_\alpha^2 - Y_\beta^2]$$

coefficient is model independent

- Resulting bounds on Majorana  $\nu$  transition moments:

Dim-5  
“matching”

$$\frac{[\mu_\nu]_{\alpha\beta}}{\mu_B} \leq 4 \times 10^{-9} \left( \frac{[m_\nu]_{\alpha\beta}}{1 \text{ eV}} \right) \left( \frac{1 \text{ TeV}}{\Lambda} \right)^2 \left| \frac{m_\tau^2}{m_\alpha^2 - m_\beta^2} \right|$$

Strong dependence on  $\Lambda$

Dim-7  
“mixing”

$$\frac{[\mu_\nu]_{\alpha\beta}}{\mu_B} \leq 1 \times 10^{-7} \left( \frac{[m_\nu]_{\alpha\beta}}{1 \text{ eV}} \right) \left( \log \frac{\Lambda}{M_Z} \right)^{-1} \left| \frac{m_\tau^2}{m_\alpha^2 - m_\beta^2} \right|$$

Mild logarithmic dependence on  $\Lambda$

At  $\Lambda \sim 1 \text{ TeV}$  both bounds are weaker than experimental limits



## Do we understand the difference of Dirac vs Majorana bounds ?

Yes.

The difference can be traced to different nature of “Voloshin’s symmetry” protecting Dirac and Majorana mass terms, and the way this symmetry is broken by SM interactions.

# Do we understand the difference of Dirac vs Majorana bounds ?

- **Voloshin's mechanism:** identify  $SU(2)_V$  under which
  - magnetic moment transforms as a singlet (allowed)
  - mass transform as a triplet (forbidden)

Two flavors

$$\psi_V = \begin{pmatrix} \nu_L \\ \nu_R^c \end{pmatrix} \quad \text{is a doublet under } SU(2)_V \text{ (Dirac)}$$
$$\psi_V = \begin{pmatrix} \nu_L^1 \\ \nu_L^2 \end{pmatrix} \quad \text{Is a doublet under } SU(2)_V \text{ (Majorana)}$$

- If underlying s.d. dynamics is invariant under  $SU(2)_V$ , then one can have large magnetic moments and small mass.

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- Naturalness bounds rely on the fact that:
  - $SU(2)_V$  (Dirac) is broken by SM **gauge interactions**
  - $SU(2)_V$  (Majorana) is broken by SM **Yukawa interactions**

# Conclusions

- Naturalness bounds on  $\mu_\nu$  obtained by requiring that mass induced by rad. corr. does not exceed 1 eV:

Dirac

$$\frac{|\mu_\nu|}{\mu_B} \lesssim 3 \times 10^{-15} \times \left( \frac{\delta m_\nu}{1 \text{ eV}} \right) \times \left( \frac{1 \text{ TeV}}{\Lambda} \right)^2$$

Majorana

$$\frac{[\mu_\nu]_{\alpha\beta}}{\mu_B} \leq 4 \times 10^{-9} \left( \frac{[m_\nu]_{\alpha\beta}}{1 \text{ eV}} \right) \left( \frac{1 \text{ TeV}}{\Lambda} \right)^2 \left| \frac{m_\tau^2}{m_\alpha^2 - m_\beta^2} \right|$$

- Dirac vs Majorana difference understood in terms of different Voloshin's symmetries
- Discovery of neutrino magnetic moment near present limits would strongly suggest that neutrinos are Majorana particles



# Additional Material

- Voloshin's symmetry for Dirac neutrinos

$$\psi = \begin{pmatrix} \nu_L \\ \nu_R^c \end{pmatrix} \quad \text{transform as a doublet under } \text{SU}(2)_V$$

Singlet	→	$\mathcal{L}_\mu \propto \mu_\nu \psi^T \left\{ [C\sigma_{\mu\nu}] \otimes \begin{pmatrix} 0 & -1 \\ 1 & 0 \end{pmatrix} \right\} \psi \cdot F^{\mu\nu}$
		$\uparrow$ $\text{SU}(2)_V$ structure $\downarrow$
Triplet	→	$\mathcal{L}_m \propto m_\nu \psi^T \left\{ [C] \otimes \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix} \right\} \psi$

- $\text{SU}(2)_V$  broken by gauge interactions  $\Rightarrow \delta m_\nu \propto \frac{\alpha}{\pi} \Lambda^2 \mu_\nu$

- Voloshin's symmetry for Majorana neutrinos (two flavors)

$$\psi = \begin{pmatrix} \nu_L^1 \\ \nu_L^2 \end{pmatrix} \quad \text{transform as a doublet under } \text{SU}(2)_V$$

Singlet

$$\mathcal{L}_\mu \propto [\mu_\nu]_{12} \psi^T \left\{ [C\sigma_{\mu\nu}] \otimes \begin{pmatrix} 0 & -1 \\ 1 & 0 \end{pmatrix} \right\} \psi \cdot F^{\mu\nu}$$

SU(2)<sub>V</sub> structure

Triplet

$$\mathcal{L}_m \propto \psi^T \left\{ [C] \otimes \begin{pmatrix} (m_\nu)_{11} & (m_\nu)_{12} \\ (m_\nu)_{12} & (m_\nu)_{22} \end{pmatrix} \right\} \psi$$

- SU(2)<sub>V</sub> broken by Yukawa interactions  $\Rightarrow$

$$\delta[m_\nu]_{\alpha\beta} \propto \frac{\alpha}{\pi} \Lambda^2 [\mu_\nu]_{\alpha\beta} (Y_\alpha^2 - Y_\beta^2)$$