GRAVITATIONAL WAVE ASTRONOMY

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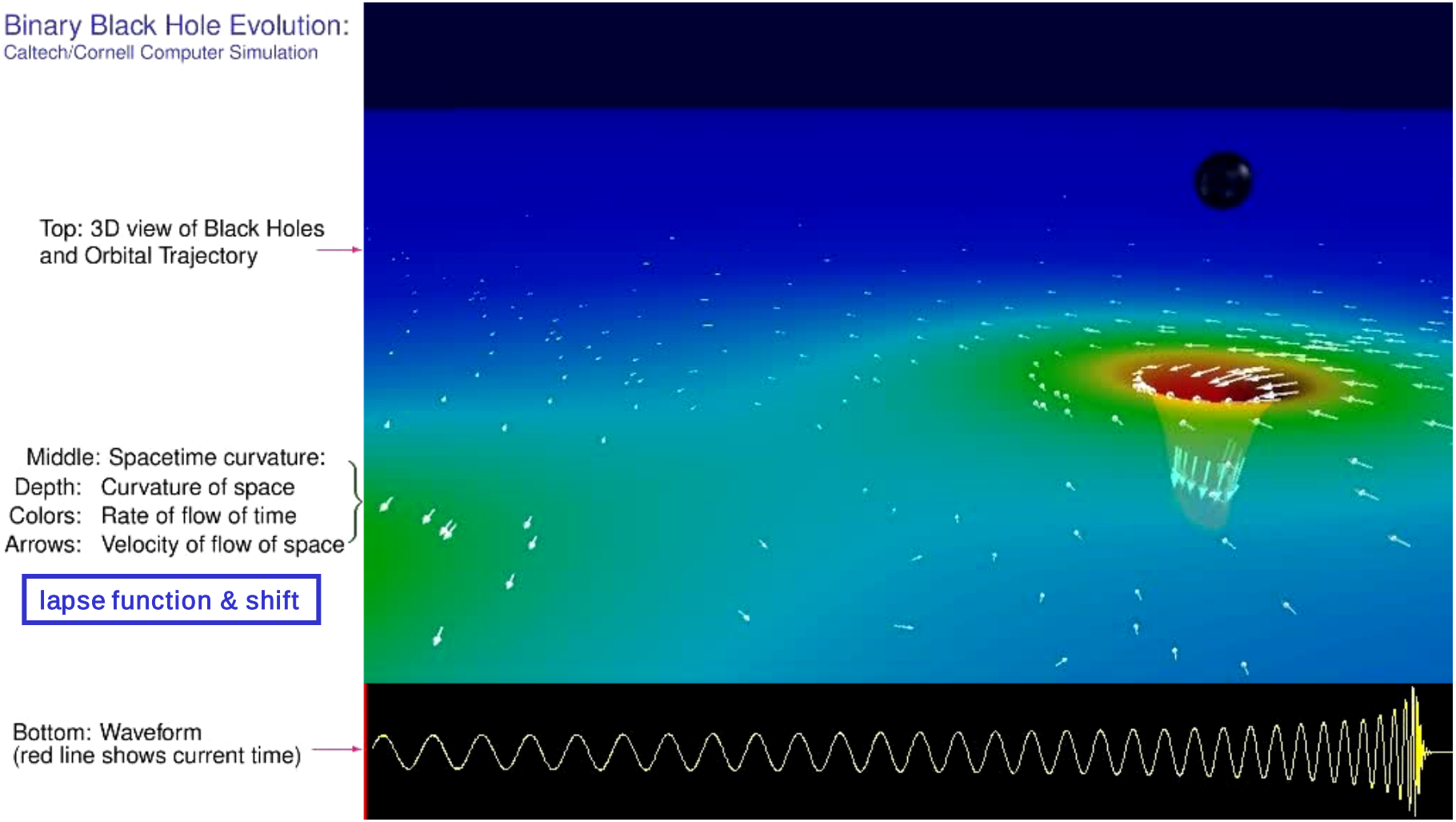
1. **GW: physics & astronomy**
2. Current- & next-gen detectors & searches
3. Burst sources: CBC, SN → GR, cosmology
4. Periodic sources: NS → subatomic physics

LECTURE ONE

- GW in pictures and sounds
- Compendium of **science questions**

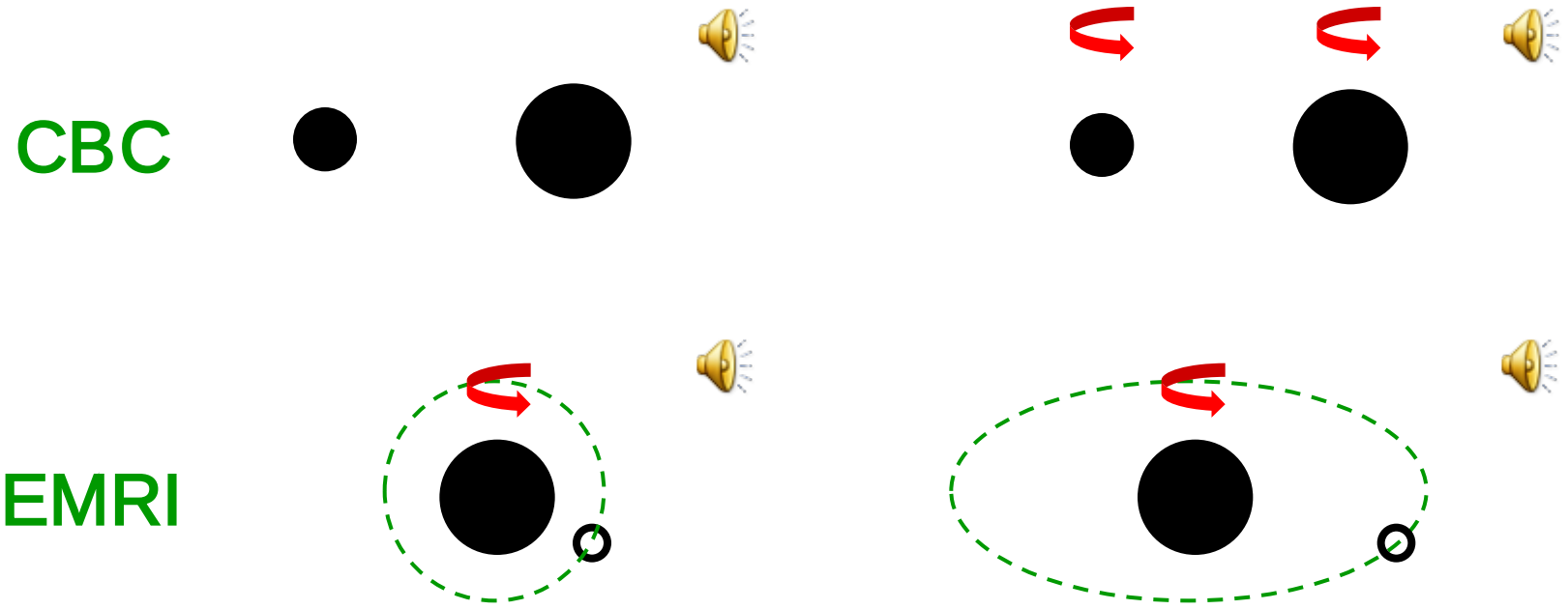
- GW physics: polarization, multipoles
- Compact binary coalescence

- The future is here… e.g. **GW cosmology**



Test strong-field GR via ring-down normal modes (Kerr?)

EINSTEIN'S MUSIC



Courtesy Scott Hughes @ gmunu.mit.edu

$h_+ \rightarrow$ chirp; spin-orbit precession; GW circularizes

By “listening” closely to various GW tones,
we can probe fundamental physics
under conditions **inaccessible on Earth**

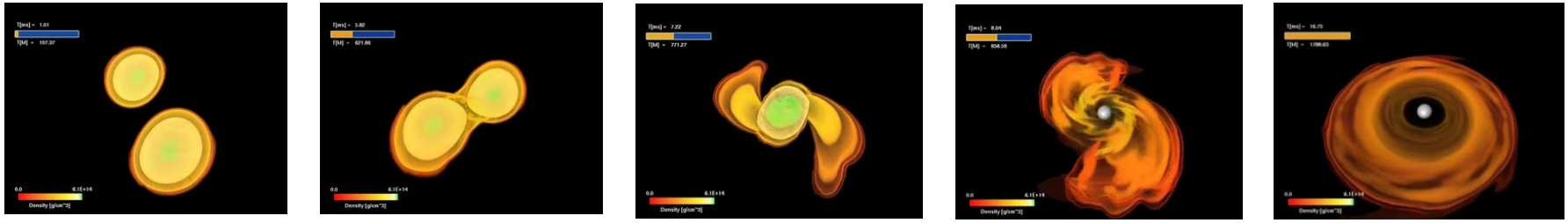
BIG QUESTIONS

- Strong-field tests of General Relativity
  → do Nature's black holes match the rotating Kerr solution?
  → alternative gravities? (speed, mass, polarization of graviton)
  → frame dragging, spin precession, ... (from merger waveforms)
  → black hole oscillations and existence of naked singularities
- Tests of Quantum Chromodynamics
  → nuclear equation of state (n radius, strange quark matter)
  → many-body phenomena (QGP, superconductivity; cf. LHC)
- Birth of the Universe
  → fluctuations at Planck scale before surface of last scattering
  → precision cosmology (e.g. H_0) with EM counterparts
  → phase transitions? (topological defects like strings)

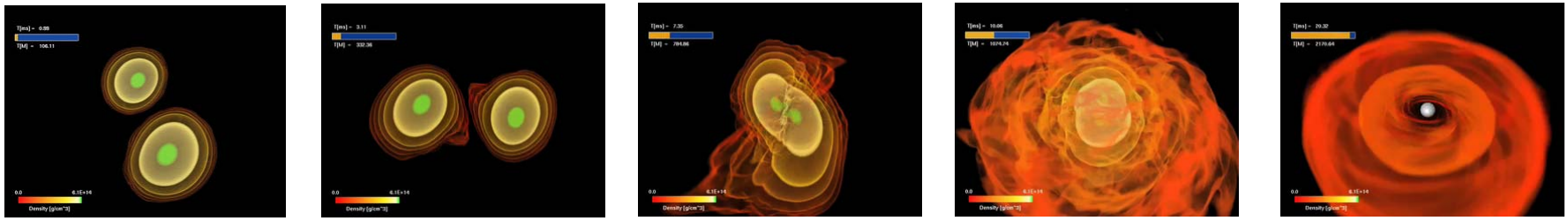
- Multi-messenger astronomy = EM + GW + ν
  - → hydro, particle, and magnetic physics in core-collapse SN
  - → what are γ-ray bursts? (brightest things in the sky...)
  - → extreme magnetism (10^{15} G), e.g. explosive annihilation
  - → how do black holes form, and how many are there?
  - → unexpected new sources!
- Macroscopic quantum metrology
  - → squeezed vacuum, quantum non-demolition measurements
  - → stability and control of opto-mechanical systems
  - → ultra-stable oscillators and lasers, ultra-smooth mirrors, AO
- Massively distributed computing

- **Cool pure-GW discoveries already!**
  - → NS deformation and internal magnetisation
  - → Kill off some superstring models

GW → NUCLEAR PHYSICS

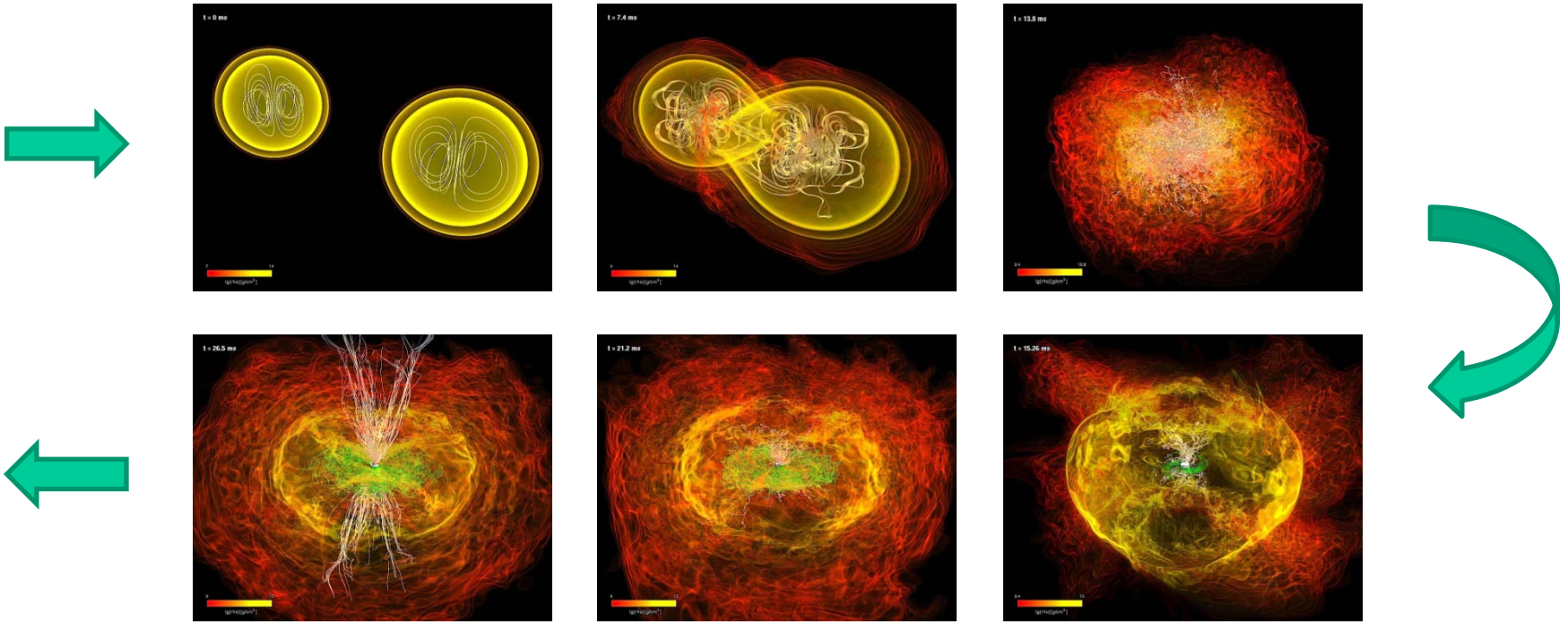


Polytropic equation of state (density contours)

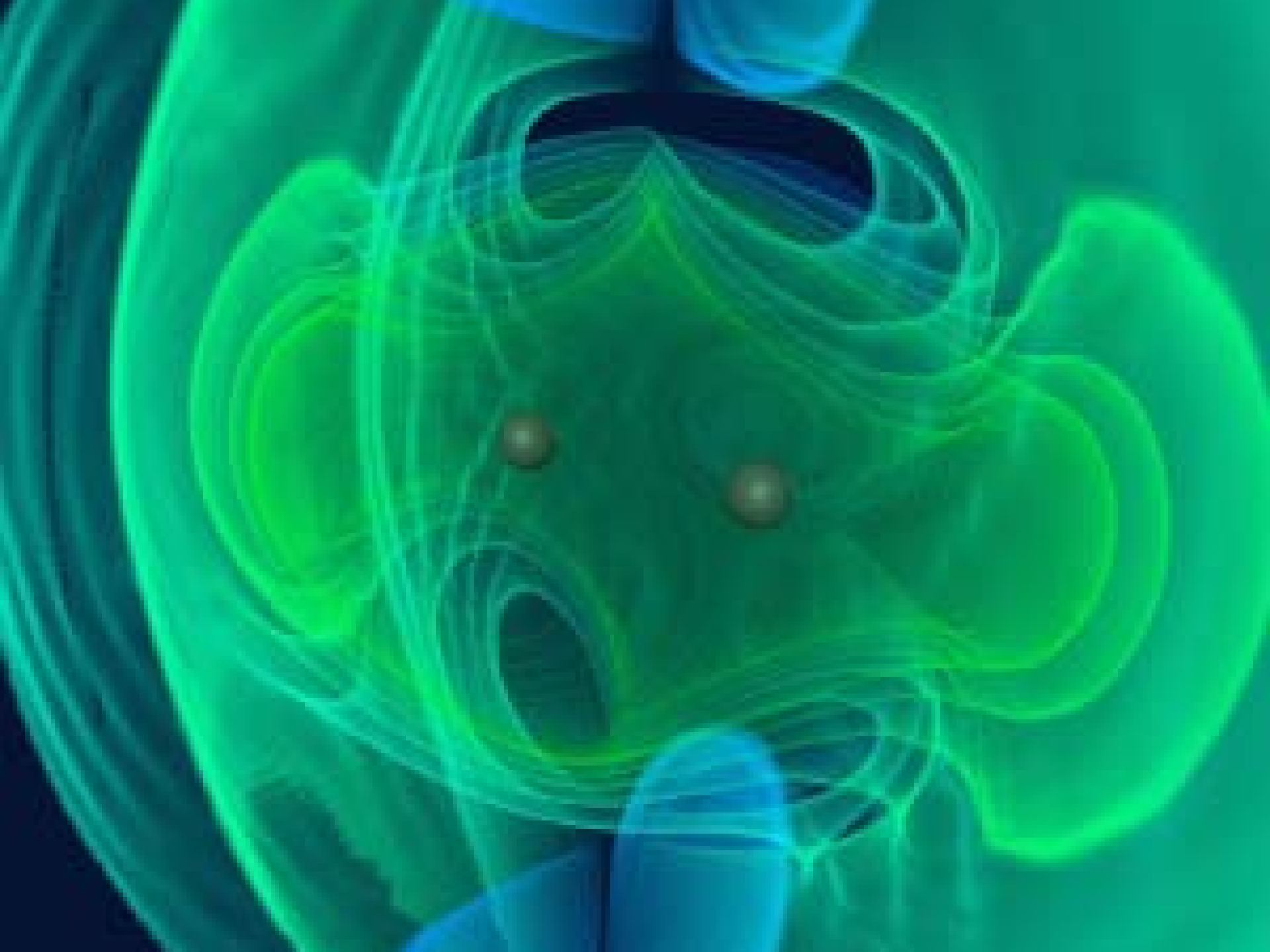


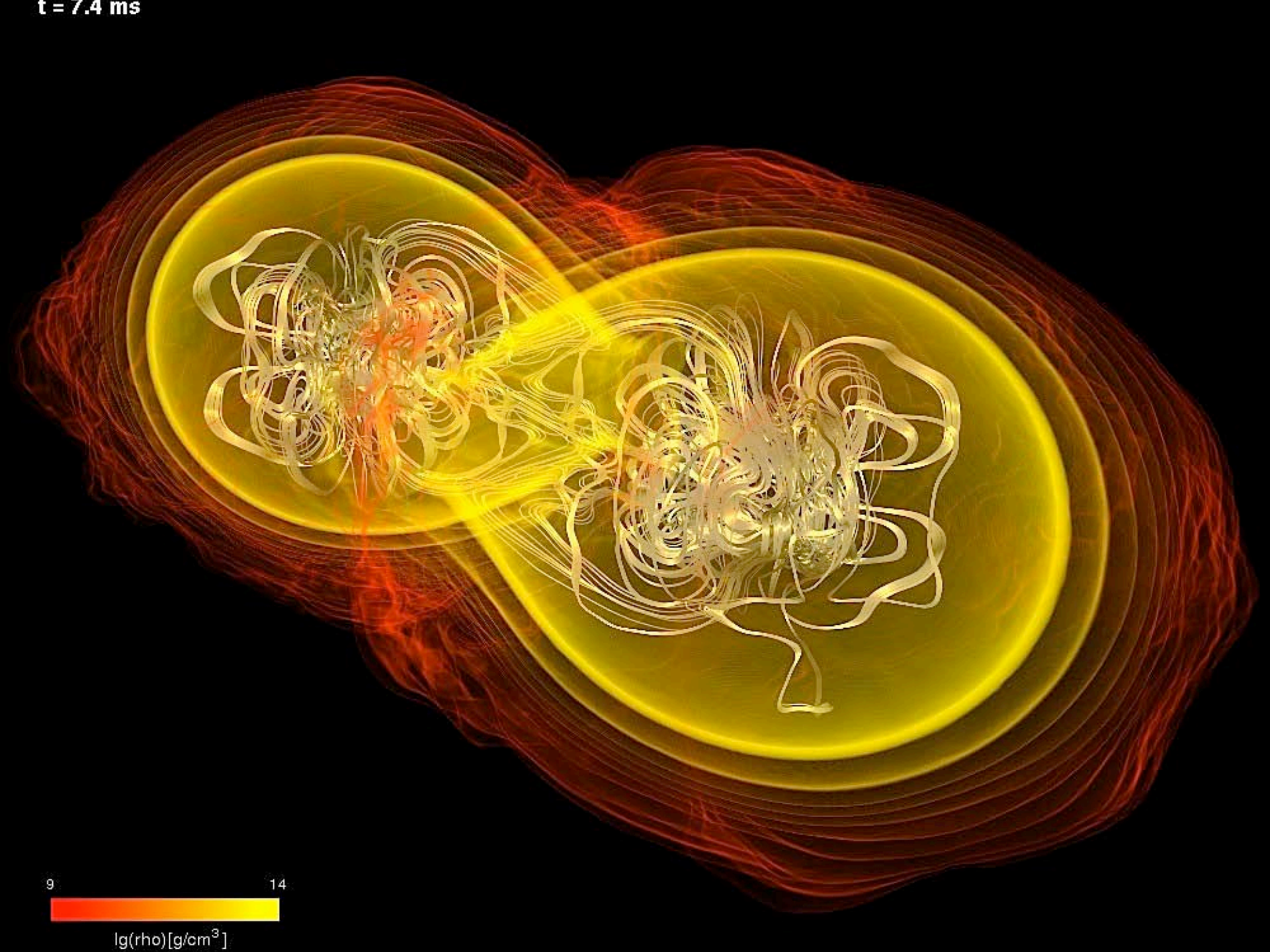
“Superfluid” equation of state

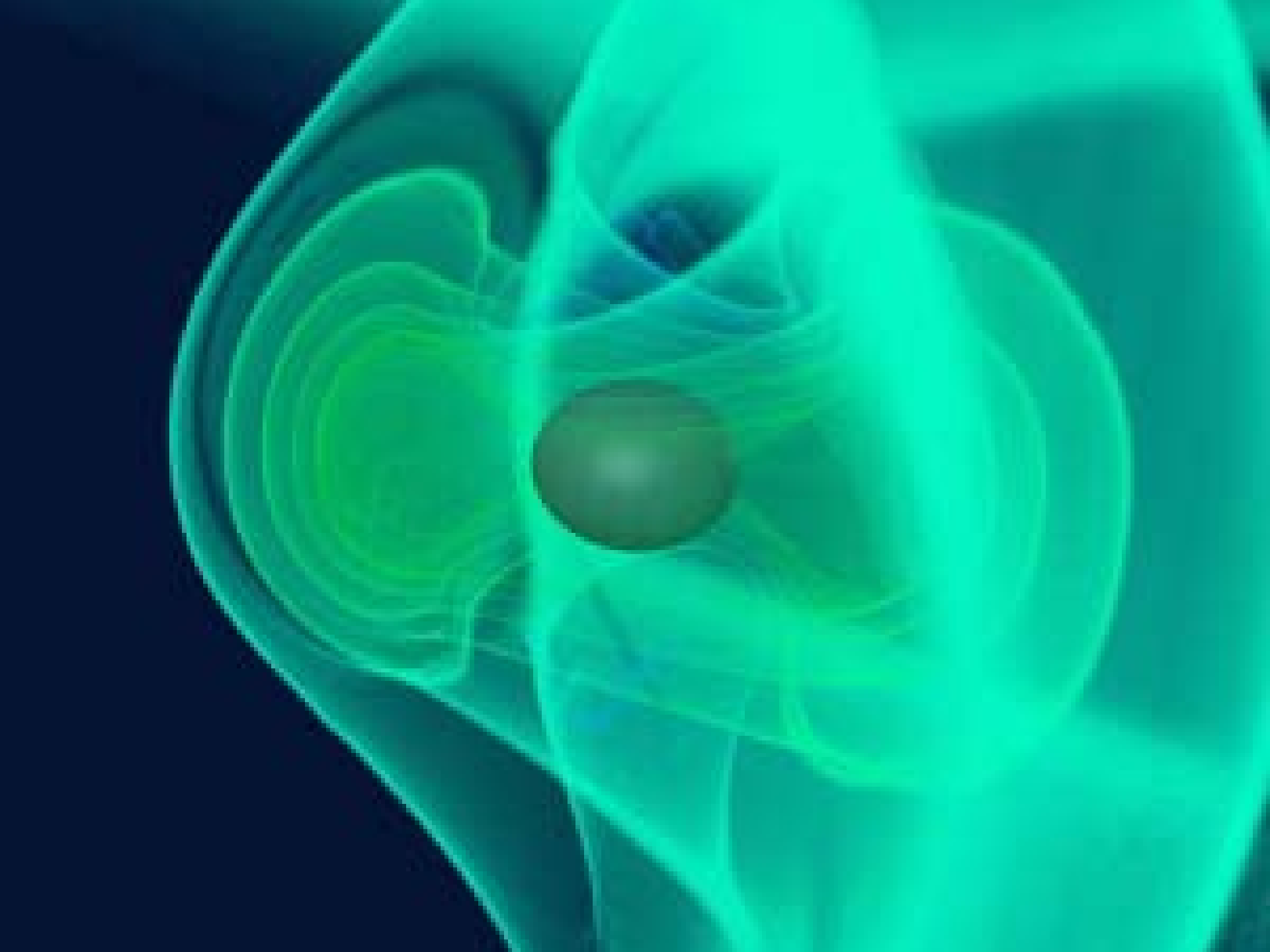
GW → ASTROPHYSICS

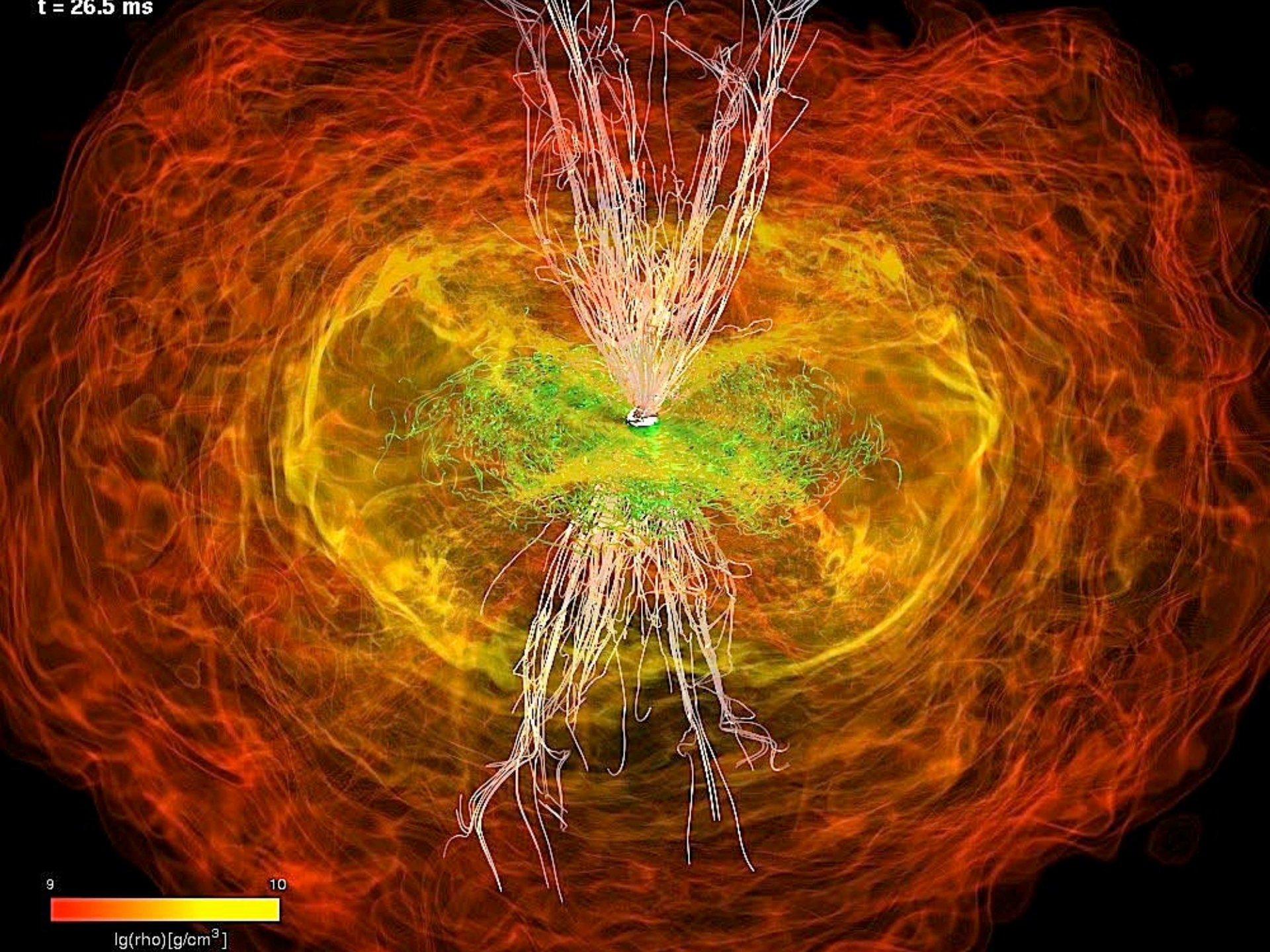


How to cook up a magnetized, relativistic jet
in a “short” (CBC) gamma-ray burst… (numrel.aei.mpg.de)




t = 7.4 ms
9
14
lg(rho)[g/cm^3]




t = 26.5 ms
9
10
lg(rho)[g/cm^3]

WHAT ARE GW?

A bit of theory…

RIPPLES IN SPACETIME

- Harmonic vibration of the metric $g_{\alpha\beta}$ and hence curvature $R_{\alpha\beta\gamma\delta}[g_{\alpha\beta}]$ in vacuum

$$R^{\mu}{}_{\alpha\mu\beta} - \tfrac{1}{2} R^{\mu\lambda}{}_{\mu\lambda} g_{\alpha\beta} = 0 = (8\pi G / c^4) T_{\alpha\beta}$$

- Propagates at c

- Interval between two events = **flat space** + **gravitational wave**

$$ds^2 = c^2 dt^2 - dx^2 + h_{\alpha\beta} dx^{\alpha} dx^{\beta}$$

TWIN POLARIZATIONS

- Bianchi → six polarizations (e.g. breathing)
- GR → $R^{\mu}{}_{\alpha\mu\beta}=0$ → transverse, traceless gauge

$$h_{ij}(z,t) = h^{+}(t\text{-}z/c)e^{+}_{ij} + h^{\times}(t\text{-}z/c)e_{ij}$$

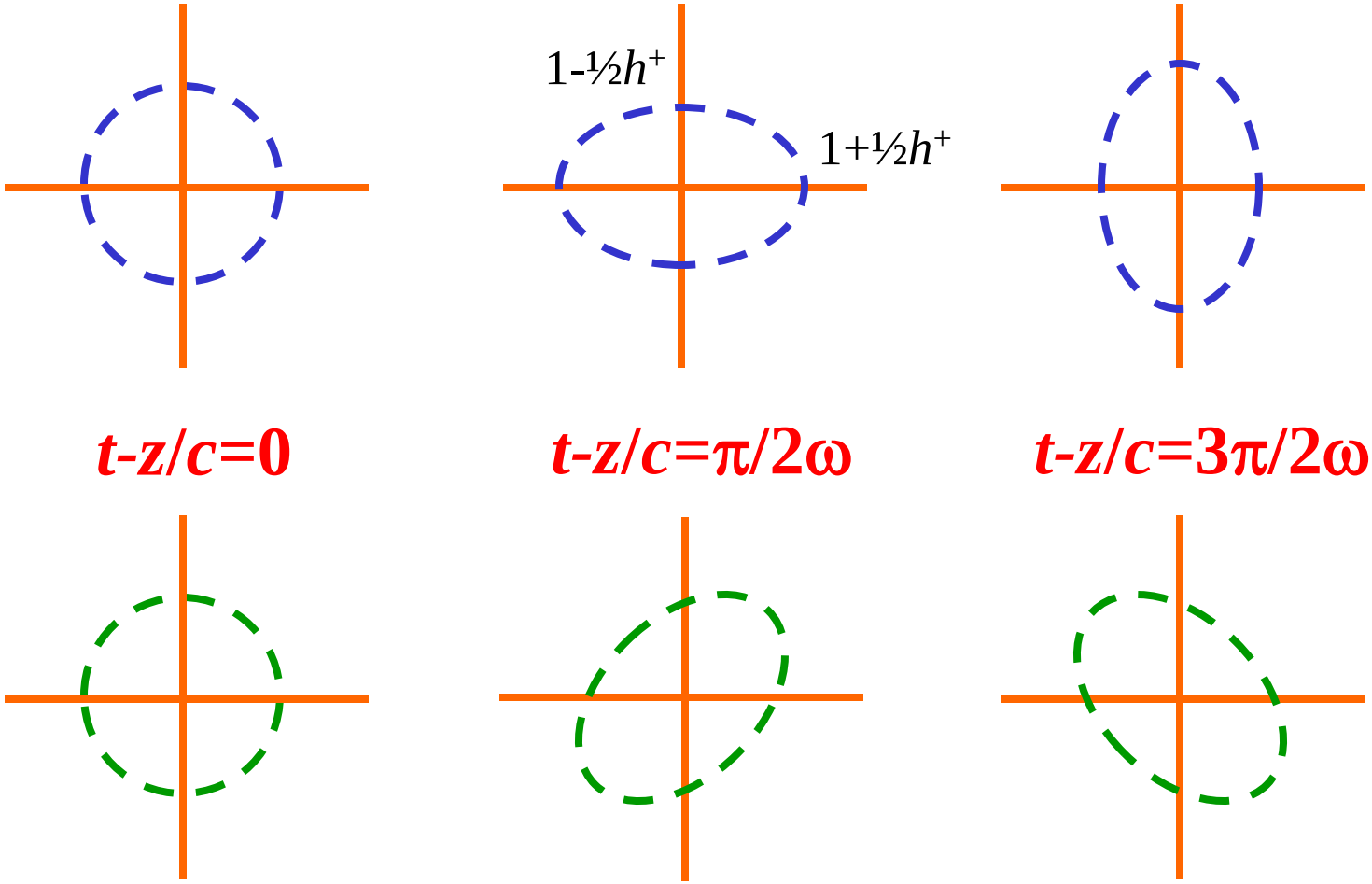
- TIDAL acceleration: no “00” component

Plus mode: $e^{+}_{xx} = 1 = -e^{+}_{yy}$

- $ds^2 = c^2dt^2 - (1-h^{+})dx^2 - (1+h^{+})dy^2 - dz^2$

Cross mode: $e^{+}_{xy} = 1 = e^{+}_{yx}$

- $ds^2 = c^2dt^2 - dx^2 - dy^2 - 2h^{\times}dx\,dy - dz^2$



Locus of events “equidistant” from (0,0)
(ring of test masses)

DIPOLE ANTENNA

- Mass dipole moment

$$q_i = \Sigma m x_i$$

- Power output $\propto (d^2 q_i / dt^2)^2$ **vanishes**
  (momentum conserved if zero external force)

$$d^2 q_i / dt^2 = \Sigma m \, d^2 x_i / dt^2 = 0$$

- Nonzero external force (“giant wrench”) → equal but opposite dipole moment

GW SOURCE PHYSICS

- Mass quadrupole $\propto$ energy density $\propto T^{00}$

$$h_{lm} = \frac{G}{c^4 d}\frac{\partial^2}{\partial t^2}\int d^3x Y_{lm}^* r^l \rho$$

$$\frac{GM}{Rc^2}\left(\frac{Rf}{c}\right)^2\left(\frac{d}{R}\right)^{-1}\varepsilon_\rho$$

- Current quadrupole $\propto$ momentum flux $\propto T^{0i}$

$$h_{lm} = \frac{G}{c^5 d}\frac{\partial^2}{\partial t^2}\int d^3x Y_{lm}^* r^l \mathbf{x}\cdot\mathrm{curl}(\rho\mathbf{v})$$

$$\frac{GM}{Rc^2}\left(\frac{Rf}{c}\right)^2\left(\frac{d}{R}\right)^{-1}\frac{\varepsilon_{\rho\mathbf{v}}V}{c}$$

SOME USEFUL FORMULAS

- **Metric perturbation** (Thorne 80 RevMP)

$$h_{jk}^{TT} = \frac{G}{c^4 d}\sum_{l=2}^{\infty}\sum_{m=-l}^{l}\left[\partial_t^{(l)} I^{lm} T_{jk}^{E2,lm} + \partial_t^{(l)} S^{lm} T_{jk}^{B2,lm}\right]$$

MASS CURRENT

- **Antenna pattern**

$$T_{jk}^{E2,lm}(\theta,\varphi) \propto r^2 [\nabla\nabla Y^{lm}]_{jk}^{STT} \qquad T_{jk}^{B2,lm}(\theta,\varphi) \propto [(\mathbf{x}\times\nabla)(r\nabla Y^{lm})]_{jk}^{STT}$$

- **Total power**

$$\frac{dE}{dt} = \frac{G}{32\pi c^4}\sum_{l=2}^{\infty}\sum_{m=-l}^{l}\left[\left|\partial_t^{(l+1)} I^{lm}\right|^2 + \left|\partial_t^{(l+1)} S^{lm}\right|^2\right]$$

COMPACT SOURCES

$h \approx (V/c)^2 (r/R_s)^{-1}$ $\qquad$ $L \approx c^5 G^{-1} (V/c)^6 (R/R_s)^{-2}$

- Hulse-Taylor binary: M_{Sun}, 10^9 m, 10^5 ms^{-1}
  $h \approx \mathbf{10^{-24}}$ @ 10 kpc $\leftrightarrow$ $L \approx 10^{27}$ erg s^{-1} **kHz**

- NS mountain: εM_{Sun}, 10^4 m, 10^7 ms^{-1}
  $h \approx \mathbf{10^{-16}\varepsilon}$ @ 1 kpc $\leftrightarrow$ $L \approx 10^{49}\varepsilon^2$ erg s^{-1} **kHz**

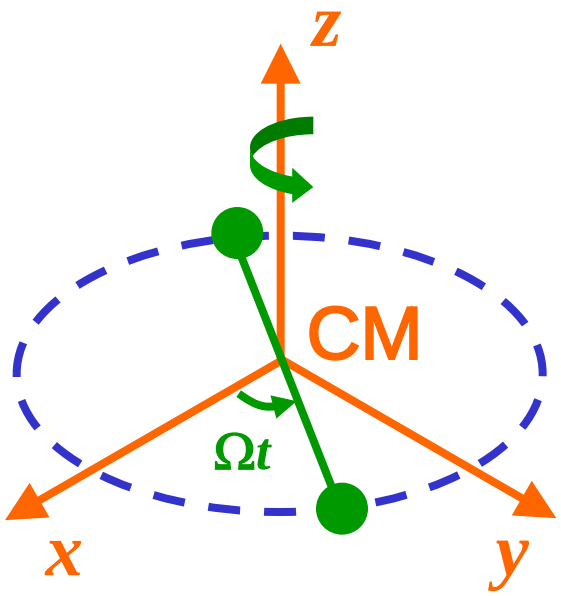
- Merging galaxies: $10^6 M_{Sun}$, 1 pc, 10^6 ms^{-1}
  $h \approx \mathbf{10^{-18}}$ @ 1 Mpc $\leftrightarrow$ $L \approx 10^{36}$ erg s^{-1} **mHz**

EX: COALESCING BINARY

- $I_{xx} = -I_{yy} = \frac{1}{2}\mu a^2 \cos(2\Omega t)$
  $I_{xy} = I_{yx} = \frac{1}{2}\mu a^2 \sin(2\Omega t)$

- Observer along z axis:
  $h^+ = \cos(2\Omega t)$, $h^\times = \sin(2\Omega t)$

- Observer along y axis:
  $h^+ = \frac{1}{2}\cos(2\Omega t)$, $h^\times = \mathbf{0}$



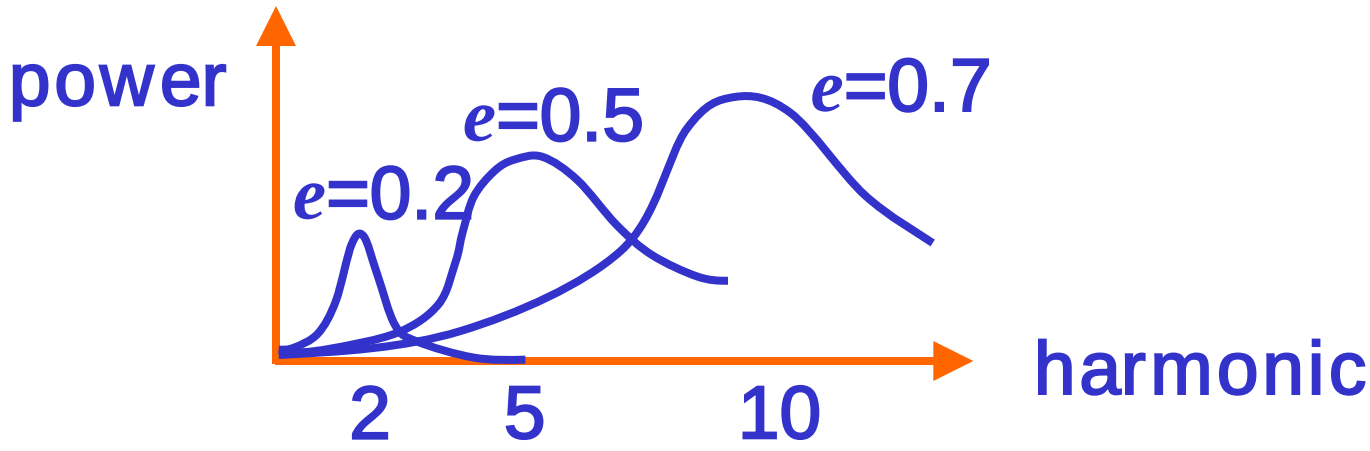
CIRCULAR ORBIT

- h_{ij} oscillates at $2\Omega_{orb}$
- Orbit **shrinks**
- Coalesces after

$$47\ \mathrm{Gyr}\ (P_{orb}/1\ \mathrm{day})^{8/3}\ (M_1/M_{Sun})^{-5/3}\ q(1+q^{-1})^{1/3}$$

ELLIPTICAL ORBIT

- **All** harmonics of Ω_{orb}
- Orbit shrinks and **circularizes**
- GW enhanced ($\times$10 for e=0.6) (Peters 73)



GW are neither scattered nor absorbed;
they map $T^{\mu\nu}$ transparently

(e.g. Thorne's "lattice" of NSs in space)

PRECISION GW COSMOLOGY

New results already!

PRIMORDIAL BACKGROUND

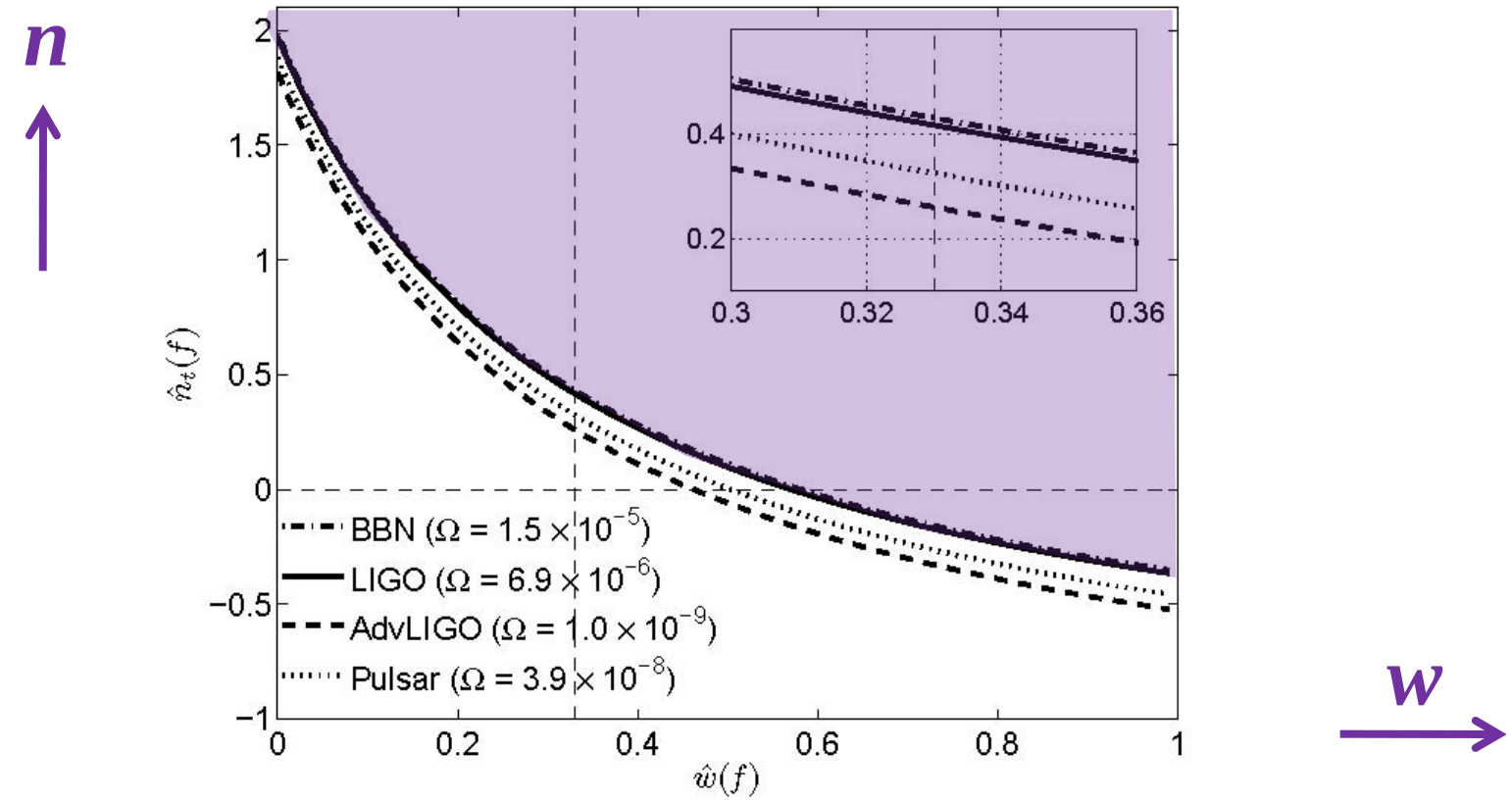
- Vacuum fluctuations amplified during **inflation** *before the surface of last scattering*
- Stochastic → cross-correlate IFOs (42-169 Hz)

$$\Omega_{GW}(f) = \frac{d(\rho_{GW}/\rho_c)}{d(\ln f)} = \Omega_\alpha (f/100\,\mathrm{Hz})^\alpha \leq 6.9\times10^{-6}$$

- Beats limits on ρ_{GW}/ρ_c from (Abbott et al. 2009)

  → **light element** abundances: $\Omega_0 \leq 1.1 \;\; 10^{-5}$

  → **CMB** spectrum: $\Omega_0 \leq 9.5 \;\; 10^{-6}$



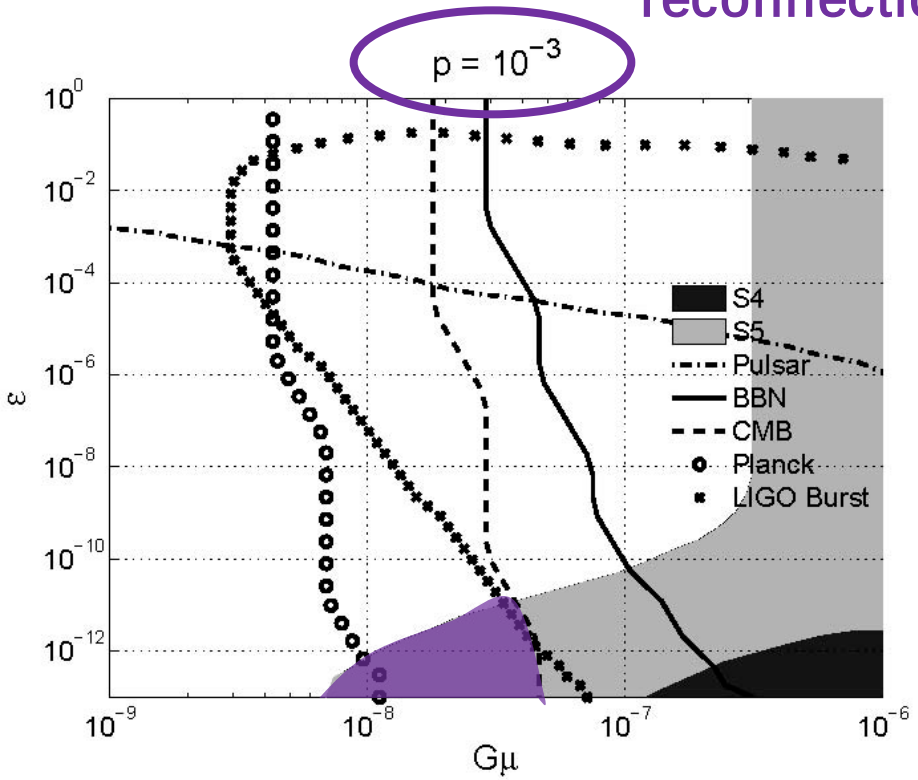
GW spectrum $\propto f^{\alpha}$ depends on $\alpha(n,w)$

$\rightarrow$ **tilt** in fluctuation spectrum $P(k) \propto k^{1+n}$

$\rightarrow$ dark energy **equation of state** $w = p/\rho$

reconnection probability

loop size



(Abbott et al. 09 Nat)

tension

Cosmic strings (reconnecting tangle)

→ radiate when cusps point at Earth

→ loop size ↔ GW radiation reaction

→ string tension ↔ vibration frequency

STANDARD SIRENS

- **Short GRB** with host galaxy **red shift** (Dalal et al. 06; Nissanke et al. 10) or without (Finn 96)
- In-spiral waveform known
- Absolute distance to surface of last scattering (angular scale of CMB peaks, matter density from peak heights)
- **Dark energy EOS:** w to 9% and H_0 to 2%
- Even at low z, **luminosity distance** $\propto \int dz\, H(z)^{-1}$ contains energy densities at all z

- GW data & Post-Newtonian waveform

  → red shifted chirp mass $(1+z)\,(m_1 m_2)^{3/5}(m_1+m_2)^{1/5}$

  → coalescence time & phase

  → **luminosity distance**

  → **sky position** & **orbital axis**

- Three “phase” parameters exquisite!
- Five “amplitude” parameters so-so…

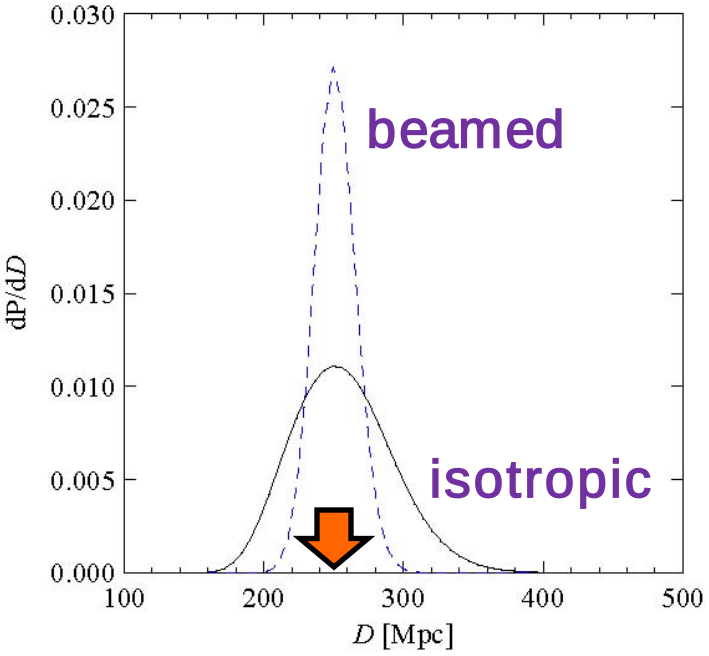
- **Distance** and **orbital inclination** degenerate

  → break with detector network (e.g. Australia)
- **Are GRBs beamed?** (Soderberg et al. 06)

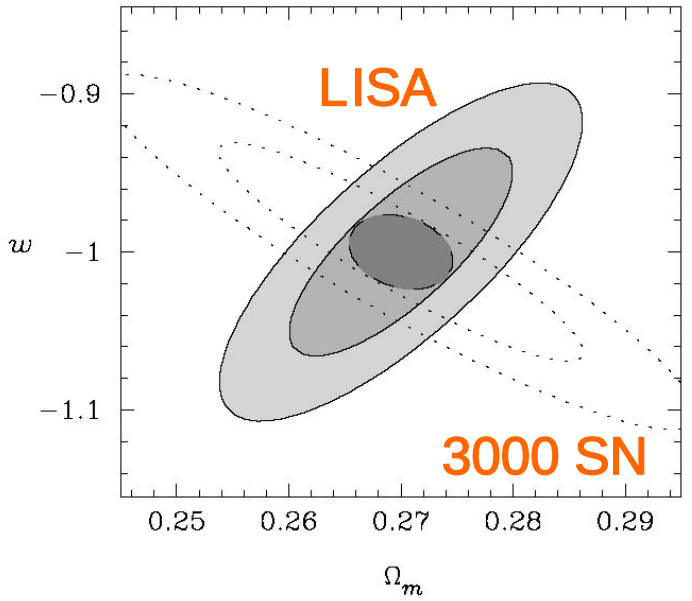
  → line of sight ∥ orbital axis

- Monte-Carlo distances for random orientations (Dalal et al. 06)
- **"No beaming"** error

  $\approx 9\ (D/0.2\mathrm{Gpc})\ \%$
- CMB errors dominate



- **LISA SMBH** in-spirals
- Distance error **$\approx 1\%$**
- **☹ Lensing fluctuations**
- EM counterparts? (Milosavljevic & Phinney 05)



SUMMARY

- GW in pictures and sounds
- Compendium of **science questions**

- GW physics: polarization, multipoles
- Compact binary coalescence

- The future is here… e.g. **GW cosmology**