

2001 July Lecture: Out of Africa

The 100th Anniversary of Enrico Fermi

**Out of Africa:
A 2 billion year old nuclear reactor**

2001 July Lecture
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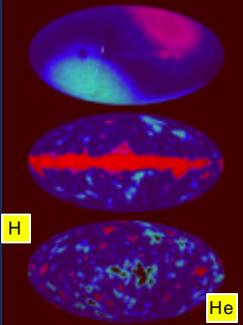
Outline

- In the beginning...
 - The Big Bang
 - Stellar nucleosynthesis
 - Formation of the Earth and solar system
- On the Nucleus, Isotopes and the Strong Nuclear Force
- Nuclear fission
- Enrico Fermi and the first artificial nuclear reactor 58 y BP
- No really, the first nuclear reactor: The Oklo phenomenon 2 Gy BP
- Implications for the future

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In the Beginning...

Echoes of the Big Bang

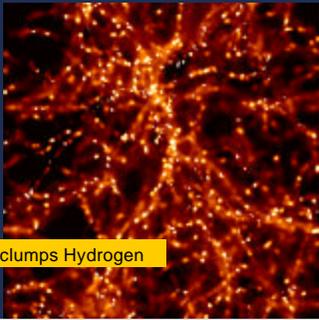


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Clumping into Stars & Galaxies

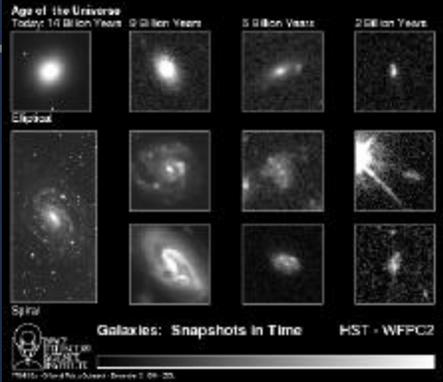


Gravity clumps Hydrogen

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Age of the Universe

Today: 14 Billion Years 9 Billion Years 5 Billion Years 2 Billion Years



Elliptical

Spiral

Galaxies: Snapshots in Time HST - WFPC2

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Nucleosynthesis builds Elements



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Star Supernovae Seed Interstellar Medium

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Radioactive elements formed in supernovae

Crab Nebula

NGC 6948

HST - WFC2

Supernovae remnants

Stellar Ashes Build New Planets

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Stars and Planets form in Orion

Earth built from stellar ashes

K

U

Th

To make nuclear reactors: need two of these!

Radioactivity

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- The stability of the nucleus depends on the number of protons and neutron in the nucleus
- Some isotopes are stable indefinitely
- Others are not.....
- Discovered by Henri Becquerel in 1896
 - He discovered Uranium ore would expose photographic film
- Marie Curie studied "pitchblende" (UO)
 - Found in 1898 it contained material that was more radioactive than the uranium itself!
 - This was the discovery of Po, Ra (decay products of Uranium)

3 Nobel Prizes

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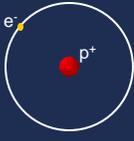
Enrico Fermi

History of neutron experiments

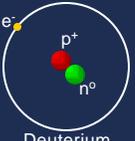
- Rn/Be neutron tubes
- The goldfish pond experiment
- Migration to the USA
- Start of WW2
- The experiments in Chicago
- The German reactor project
- (The Manhattan project)

Isotopes

- An atom consists of a nucleus surrounded by a cloud of electrons
- The nucleus consists of a bunch of protons (+ charge) and neutrons (no charge) bound together
- The chemical characteristics of the atom are determined by the number of protons in the nucleus
- An atom can have a number of different isotopes that differ only in the number of neutrons

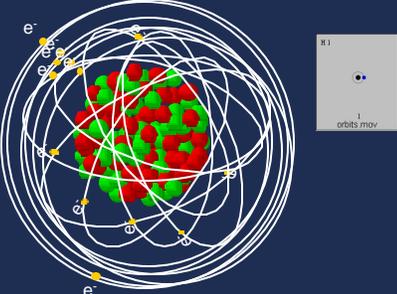


Hydrogen



Deuterium
(0.015% of all H)

Heavy atom (artist's impression)



Dust mote in a cathedral

Isotope names

Number of nucleons (protons + neutrons)

Chemical symbol (defines number of protons, 19 in this case)

40K

Data:

- Of all the potassium atoms in the Earth:
 - 93.1% are ³⁹K (stable)
 - 0.01% are ⁴⁰K (unstable, half life 1.3 billion years)
 - 6.9% are ⁴¹K (stable)

Isotope names

Number of nucleons (protons + neutrons)

Chemical symbol (defines number of protons, 6 in this case)

12C

Data:

- Of all the carbon atoms in the biosphere of Earth:
 - 98.89% are ¹²C (stable)
 - 1.11% are ¹³C (stable)
 - 1.5x10⁻¹⁰% are ¹⁴C (unstable, half life 5730 years)
- (except petrol - no ¹⁴C)

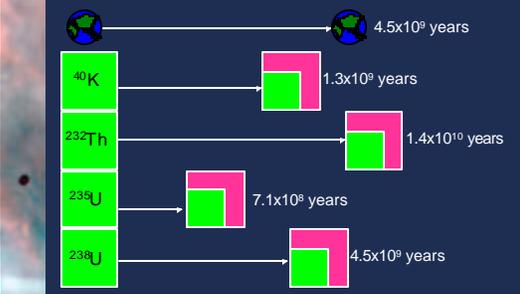
Natural Radioactive elements

H																	He
Li	Be											B	C	N	O	F	Ne
Na	Mg											Al	Si	P	S	Cl	Ar
K	Ca	Sc	Ti	V	Cr	Mn	Fe	Co	Ni	Cu	Zn	Ga	Ge	As	Se	Br	Kr
Rb	Sr	Y	Zr	Nb	Mo	Tc	Ru	Rh	Pd	Ag	Cd	In	Sn	Sb	Te	I	Xe
Cs	Ba	La	Hf	Ta	W	Re	Os	Ir	Pt	Au	Hg	Tl	Pb	Bi	Po	At	Rn
Fr	Ra	Ac	Ce	Pr	Nd	Pm	Sm	Eu	Gd	Tb	Dy	Ho	Er	Tm	Yb	Lu	
			Th	Pa	U	Np	Pu	Am	Cm	Bk	Cf	Es	Fm	Md	No	Lr	

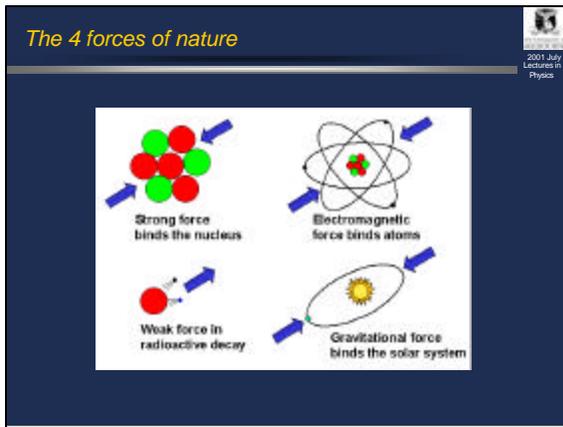
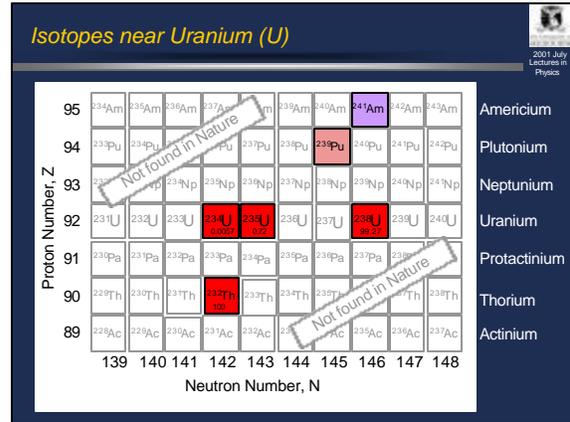
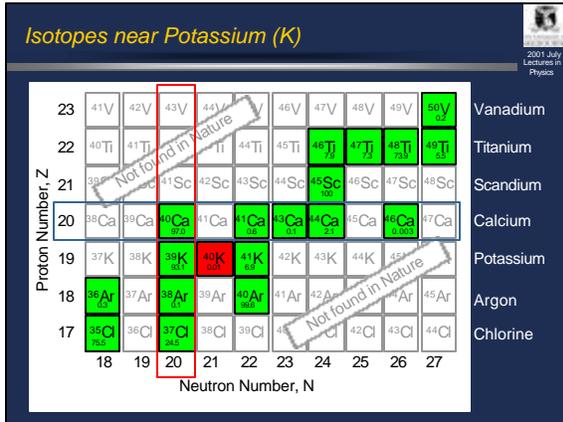
U Radioactive elements
Tc Elements not presently found in nature

Radioactive elements decay (slowly)

Radioactivity provides dates over geological ages



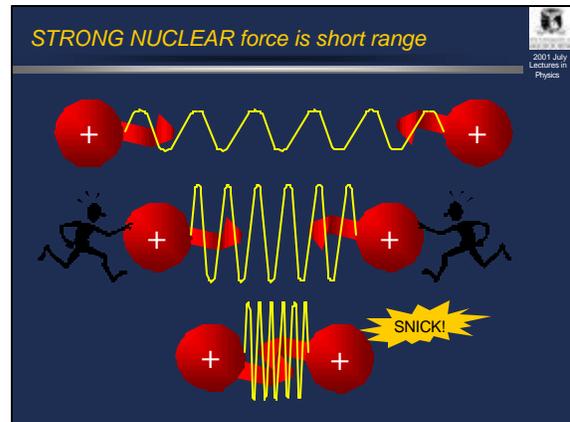
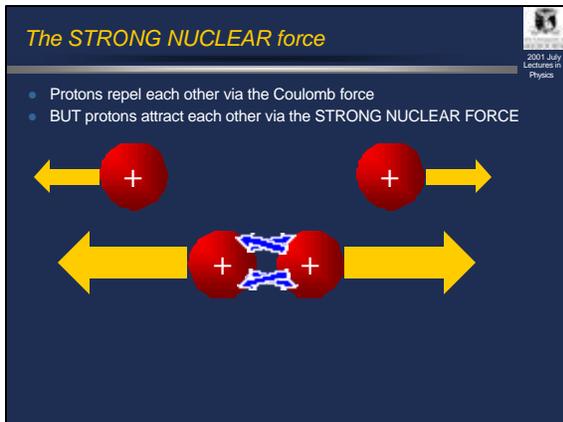
- ⁴⁰K → ⁴⁰Ar: 4.5x10⁹ years
- ²³²Th → ²⁰⁸Pb: 1.3x10¹⁰ years
- ²³⁵U → ²⁰⁷Pb: 7.1x10⁸ years
- ²³⁸U → ²⁰⁶Pb: 4.5x10⁹ years



Enrico Fermi and the path to artificial fission

- 1901: Enrico Fermi born in Rome
- 1922: Fascist government in Italy
- 1923: Contributes guest article "Mass in Relativity Theory" to an Italian translation of a German text book
 - discusses possibility of using $E = mc^2$ to release nuclear energy

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STRONG force also works for neutrons and protons

neutron proton
neutron neutron
neutron proton

$1 \times 10^{-15} \text{ m}$
 $1 \times 10^{-15} \text{ m}$

Balance of STRONG and Coulomb forces...

...determines composition of the Universe!

- Mainly H, He and precious little else!
- Mass number is number of protons and neutrons in the nucleus
- H = 1, He = 4, Fe=56, U=235
- Note log scale!
- H more than 1000 times more abundant than C

Nuclear Binding Energy

Free nucleons
Same nucleons assembled into a nucleus
 ^{56}Fe
 $E = mc^2$

Nuclear Binding Energy

^{235}U
 $E = mc^2$

Stability of the nucleus

- Most stable configuration of protons and neutrons is ^{56}Fe

Binding Energy/nucleon (MeV)
Mass Number, A
Energy released by fusion
Energy released by fission

Enrico Fermi and the path to artificial fission

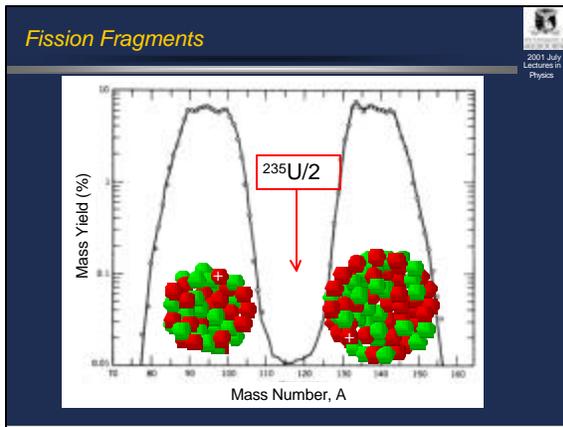
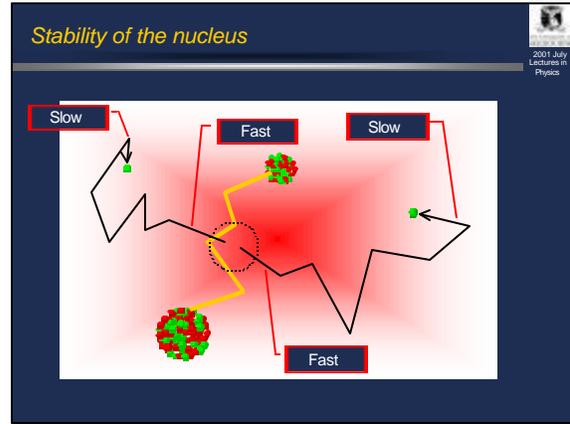
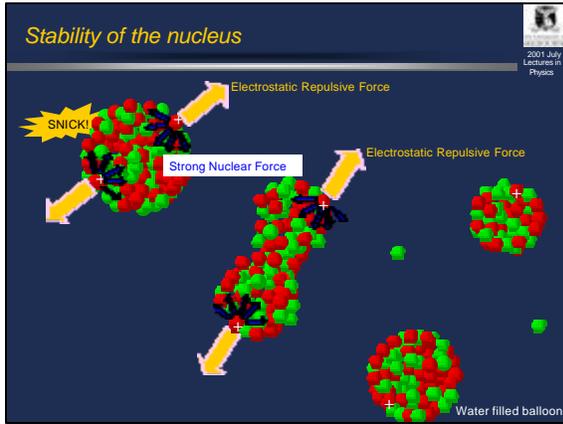
- 1933: Irene and Pierre Juliot-Curie discover that alpha particle bombardment of boron and aluminium produces new elements
- 1933: Fermi suggests trying neutron bombardment instead
 - Produces neutrons from radon gas in Be filled tubes.
 - Works his way through the periodic table.
- 1934: Fermi bombards uranium with neutrons

Juliot-Curie²
 $^{10}_5\text{B} + ^4_2\text{He} \rightarrow ^{13}_7\text{N} + ^1_0\text{n}$
 $^{27}_{13}\text{Al} + ^4_2\text{He} \rightarrow ^{30}_{15}\text{P} + ^1_0\text{n}$

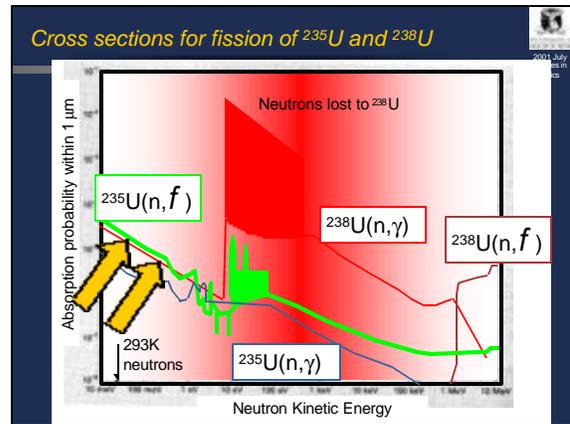
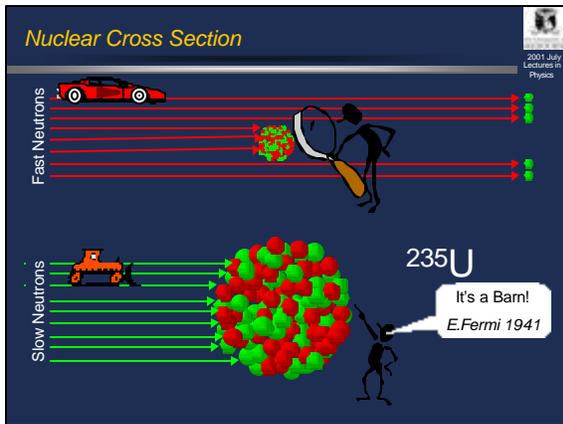
Fermi
 $^A_Z\text{X} + ^1_0\text{n} \rightarrow ^{A+1}_Z\text{X} \rightarrow ?$

Pierre Juliot-Curie patents process in Australia! 🇫🇷 🇮🇹 🇮🇹

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- ### Neutron Induced Fission
- Probability a neutron induces fission increases as the energy decreases
 - ^{235}U (rare isotope) + n leads to fission
 - ^{238}U (common isotope) + n leads to Pu
 - For a chain reaction
 - need to not lose neutrons
 - minimal neutron absorbers
 - compact body
 - need as much ^{235}U as possible
 - high probability a neutron will reach another ^{235}U nucleus
 - need to slow neutrons
 - slow neutrons have the best chance of fission
 - fast neutrons are absorbed by other processes
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1934: The failure to discover fission

- Fermi's mentor, Corbino, announces that neutron bombardment of U produces even heavier elements, Fermi skeptical
- Press announces discovery of element 93!
 - Government trumpets triumph of fascist inspired research
 - Tabloids announce jar of fascist element 93 presented to queen (fable)

$$\cancel{{}^{238}_{92}\text{U} + {}^1_0\text{n} \rightarrow \cancel{{}^{239}_{92}\text{U}} \rightarrow \cancel{{}^{239}_{93}\text{Mu}} + \text{b}^- + \bar{\text{n}}}$$

Radon+ Beryllium neutron source

1934: The failure to discover fission

- 1934 June: Nature paper "Possible production of elements of atomic number higher than 92"
- 1934 September: Ida Noddack (chemist, discoverer of Rhenium) criticises Fermi's paper - should consider the production of lighter elements
- Fermi recalculates, uses wrong mass* for He - dismisses Noddack's criticisms.
- (*not his fault)
- c.i.f. "con intuito formidabile"

The discovery of slow neutrons

- 1934: Wooden tables give better results than marble tables!
- October 1934: Discovers a paraffin filter greatly amplifies the effect
- Water also works
 - the goldfish pond experiment
- Use of slow neutrons for transmutation patented
- 1938 Dec 10: Nobel prize for slow neutron work
 - departure for USA after ceremony

Radon+ Beryllium neutron source

The failure to discover fission

- 1938 Dec 22: Fission of uranium by neutrons discovered by Otto Hahn, Fritz Strassman (in Germany) and Lise Meitner (Strassman's aunt - in exile in Sweden)
 - eclipse of German science

Hahn
Meitner

The rare isotope ${}^{235}\text{U}$ and fission

- 1939 January: "A little bomb like that (cupping hands) and it would all disappear" (Fermi, referring to Manhattan visible out the window)
- 1939 January: Leo Szilard saw the future better than anyone - tried to persuade Fermi and others to suppress uranium work
- "Nuts" E. Fermi to Szilard 1939
- 1939 February: Bohr and Fermi deduce role of ${}^{235}\text{U}$
 - ${}^{238}\text{U} + \text{n}$ leads to an odd-numbered neutron isotope
 - ${}^{235}\text{U} + \text{n}$ leads to an even numbered neutron isotope
 - Neutrons like to be in pairs
 - making ${}^{235}\text{U} + \text{n}$ liberates binding energy sufficient to fission the nucleus
 - Barrier to fission is about 6 MeV
 - Adding 1 neutron liberates about 5.3 MeV
 - Adding 1 neutron and pairing it liberates about 6.3 MeV

Einstein and the chain reaction

- 1939 July: Szilard tells Einstein about possibility of nuclear chain reaction:
 - "Daran habe ich gar nicht gedacht!" I never thought of that! A. Einstein.
- 1941 December: Fermi joins "Metallurgical laboratory", U. Chicago
- 1941 April: First carbon moderated pile built at Columbia
- "So the physicists on the 7th floor of Pupin Laboratories started looking like coal miners and the wives to whom these physicists came back tired each night were wondering what was happening" E. Fermi
- 1942 December: First reactor goes critical

Fermi's Reactor

U, UO spheres

Graphite moderator

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Other moderators

- Properties of materials used as moderators

Material	Absorption cross section for thermal neutrons	Average slowing down length from fast to thermal (cm)	
H ₂ O	0.664	5.3	Oklo
D ₂ O	0.001	11.2	CANDU
C	0.0045	19.1	Fermi

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U, UO spheres

Graphite moderator

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Approaching criticality

Layer number

- Layer 57 caused pile to "go critical" - reaction became self sustaining
- Elaborate precautions taken to prevent run-away

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Neutron "generations" in a critical reactor

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Figure 13.34 Schematic representation of processes occurring during a single generation of neutrons. The cycle has been drawn for a reproduction factor k of exactly 1.000.

Out of Africa

- The very rich uranium mines in Gabon
- Assays of uranium ore shipped to France in 1972
- Anomalous depletion of ²³⁵U: 0.4% instead of 0.72%
- Closer investigation discovered fossil nuclear reactors

Oklo: Location map

Oklo: Location picture

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From Robert Loss' Oklo website: <http://www.curtin.edu.au/curtin/centre/waisiro/OKLO>

Concentration by rivers

Life evolves and fills atmosphere with oxygen

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Sudden rise in solubility increases concentration

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Out of Africa

Peculiarities of the Oklo mine

- Incredibly rich uranium ore
- Ore deposits formed over 1.8 billion years ago
- Ore bearing strata suffered an uplift causing cracks
- Water entered cracks and penetrated ore body

From Robert Loss' Oklo website: <http://www.curtin.edu.au/curtin/centre/waisrc/OKLO>

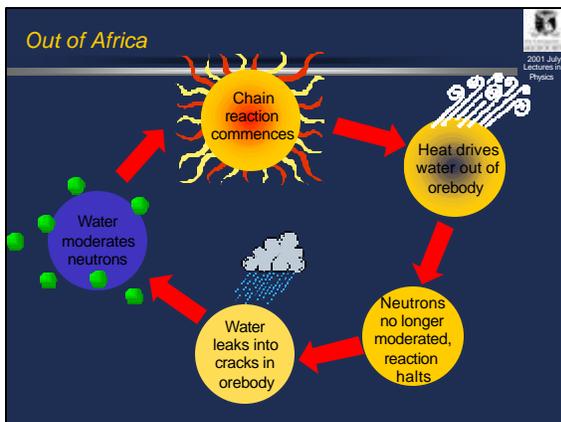
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Out of Africa

The natural reactors

- Water entering cracks in the ore body top moderated neutrons from spontaneous fission
- Silicate veins dissolved and removed allowing more water to enter
- The fission zone moved downwards through the orebody

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Natural Reactors

- 17 reactors discovered in one mine
- Several additional in a second mine nearby
- None in Brazil

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Inside the natural reactors

From Robert Loss' Oklo website: <http://www.curtin.edu.au/curtin/centre/waisrc/OKLO>

2 Billion years of weathering

Adapted from Robert Loss' Oklo website: <http://www.curtin.edu.au/curtin/centre/waisrc/OKLO>

Out of Africa

- Cycle continues until ^{235}U consumed

Conditions for the Oklo Phenomenon

Formation of the Earth (0 BY) and Now (5.0 BY) are marked on the x-axis.

Conclusion

Fermi's "pile" (as in heap):

- 6 tonnes of natural uranium (0.72% ^{235}U)
- produced 200 W of power
- used graphite as the moderator

Oklo reactors:

- generated 15,000 GW/years (large nuclear plant for 4 years)
- produced 100 kW of power
- operated for 150,000 years
- consumed 6 tonnes of uranium (3% ^{235}U)
- Used water as the moderator
- Laws of Physics same then as now
- Retained many waste products (including Pu) in place for 2 Gy

The Fermi II nuclear plant, near Detroit, USA: 1.14GW

References

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